

Generalized definition of time delay in scattering theory

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Abstract

We advocate for the systematic use of a symmetrized definition of time delay in scattering theory. In two-body scattering processes, we show that the symmetrized time delay exists for arbitrary dilated spatial regions symmetric with respect to the origin. It is equal to the usual time delay plus a new contribution, which vanishes in the case of spherical spatial regions. We also prove that the symmetrized time delay is invariant under an appropriate mapping of time reversal. These results are also discussed in the context of classical scattering theory.

1 Introduction

This paper is devoted to the definition of *time delay* (in terms of sojourn times) in two-body scattering theory. Its purpose is to advocate for the use of a *symmetrized definition* of time delay. Our main arguments supporting this point of view are the following:

- (A) In two-body scattering processes, symmetrized time delay does exist for arbitrary dilated spatial regions symmetric with respect to the origin (usual time delay does exist only for spherical spatial regions [12]). It is equal to the usual time delay plus a new contribution, which vanishes in the case of spherical spatial regions.
- (B) Symmetrized time delay is invariant under an appropriate mapping of time reversal. Usual time delay is not.
- (C) Symmetrized time delay generalizes to multichannel-type scattering processes. Usual time delay does not.

Let us recall the usual definition of time delay for a two-body scattering process in \mathbb{R}^d , $d \geq 1$. Consider a bounded open set Σ in \mathbb{R}^d containing the origin and the dilated spatial regions $\Sigma_r := \{rx \mid x \in \Sigma\}$, $r > 0$. Let $H_0 := -\frac{1}{2}\Delta$ be the kinetic energy operator in $\mathcal{H} := L^2(\mathbb{R}^d)$ and let H be a selfadjoint perturbation of H_0 such that the wave

operators $W^\pm := s\text{-}\lim_{t \rightarrow \pm\infty} e^{itH} e^{-itH_0}$ exist and are complete (so that the scattering operator $S := (W^+)^* W^-$ is unitary). Then one defines for some states $\varphi \in \mathcal{H}$ and $r > 0$ two sojourn times, namely:

$$T_r^0(\varphi) := \int_{-\infty}^{\infty} dt \int_{x \in \Sigma_r} d^d x |(e^{-itH_0} \varphi)(x)|^2$$

and

$$T_r(\varphi) := \int_{-\infty}^{\infty} dt \int_{x \in \Sigma_r} d^d x |(e^{-itH} W^- \varphi)(x)|^2.$$

If the state φ is normalized the first number is interpreted as the time spent by the freely evolving state $e^{-itH_0} \varphi$ inside the set Σ_r , whereas the second one is interpreted as the time spent by the associated scattering state $e^{-itH} W^- \varphi$ within the same region. The (usual) time delay of the scattering process with incoming state φ for Σ_r is defined as

$$\tau_r^{\text{in}}(\varphi) := T_r(\varphi) - T_r^0(\varphi). \quad (1.1)$$

For a suitable initial state φ , a sufficiently short-ranged interaction and Σ *spherical*, the limit of $\tau_r^{\text{in}}(\varphi)$ as $r \rightarrow \infty$ exists and is equal to the expectation value in the state φ of the Eisenbud-Wigner time delay operator [2, 3].

However it is shown in [12] that this limit *does not exist* if Σ is not spherical. There are other situations where the usual definition of time delay is inappropriate. Examples are multichannel-type scattering processes such as N -body scattering [13, 6, 10], scattering with dissipative interactions [9], step potential scattering [4] and scattering in waveguides [14]. In such cases time delay of the form (1.1) does not admit a limit due to the “non conservation” of the kinetic energy. Therefore one has to modify the definition (1.1) by replacing the free sojourn time $T_r^0(\varphi)$ with the *symmetrized free sojourn time* $\frac{1}{2} [T_r^0(\varphi) + T_r^0(S\varphi)]$ (see e.g. [10, Sec. V.(a)] or [14, Sec. 1]). The associated symmetrized time delay for Σ_r takes the form:

$$\tau_r(\varphi) := T_r(\varphi) - \frac{1}{2} [T_r^0(\varphi) + T_r^0(S\varphi)]. \quad (1.2)$$

Our purpose in this paper is to show that the limit of $\tau_r(\varphi)$ as $r \rightarrow \infty$ exists if Σ is an arbitrary region symmetric w.r.t. the origin and to derive an appropriate stationary formula for it.

Let us finally give a more precise description of this paper. In section 2 we introduce a natural condition on the set Σ (see Assumption 2.1) under which the existence of symmetrized time delay will be proved. We also give some auxiliary results on averaged characteristic functions. Section 3 is devoted to symmetrized time delay in classical scattering; its existence under Assumption 2.1 and its invariance under an appropriate mapping of time reversal is shown. In Section 4 we prove the same results in quantum scattering and we derive a stationary formula for the symmetrized time delay.

2 Averaged characteristic functions

Let Σ be a bounded open set in \mathbb{R}^d containing 0. For each $r > 0$ we set $\Sigma_r := \{rx \mid x \in \Sigma\}$. We shall simply say that Σ is *star-shaped* (resp. *symmetric*) whenever Σ is star-shaped (resp. symmetric) with respect to 0. Clearly Σ is star-shaped iff $\Sigma_{r_1} \subset \Sigma_{r_2}$ for $0 < r_1 \leq r_2$. Moreover to each open star-shaped set Σ we can associate a strictly positive continuous function ℓ_Σ on \mathbb{S}^{d-1} defined by

$$\ell_\Sigma(\omega) := \sup\{\mu \geq 0 \mid \mu\omega \in \Sigma\}. \quad (2.3)$$

Conversely to each strictly positive continuous function ℓ on \mathbb{S}^{d-1} one can associate a unique open star-shaped set Σ such that $\ell = \ell_\Sigma$.

We shall also consider the following class of spatial regions Σ ($\mathbb{1}_\Sigma$ stands for the characteristic function for Σ):

Assumption 2.1. Σ is a bounded open set in \mathbb{R}^d containing 0 and satisfying the condition

$$\int_0^{+\infty} d\mu [\mathbb{1}_\Sigma(\mu x) - \mathbb{1}_\Sigma(-\mu x)] = 0, \quad \forall x \in \mathbb{R}^d.$$

If $p \in \mathbb{R}^d$, then the number $\int_0^{+\infty} dt \mathbb{1}_\Sigma(tp)$ is the sojourn time in Σ of a free classical particle moving along the trajectory $t \mapsto x(t) := tp, t \geq 0$. Clearly if $\Sigma = -\Sigma$ (i.e. if Σ is symmetric), then Σ satisfies Assumption 2.1. Moreover if Σ is star-shaped and satisfies Assumption 2.1, then $\Sigma = -\Sigma$.

Lemma 2.2. Let Σ be a bounded open set in \mathbb{R}^d containing 0. Then

(a) The limit

$$R_\Sigma(x) := \lim_{\varepsilon \searrow 0} \left(\int_\varepsilon^{+\infty} \frac{d\mu}{\mu} \mathbb{1}_\Sigma(\mu x) + \ln \varepsilon \right)$$

exists for each $x \in \mathbb{R}^d \setminus \{0\}$.

(b) The (even) function $G_\Sigma : \mathbb{R}^d \setminus \{0\} \rightarrow \mathbb{R}$ given by

$$G_\Sigma(x) := \frac{1}{2} [R_\Sigma(x) + R_\Sigma(-x)]$$

satisfies

$$G_\Sigma(x) = G_\Sigma\left(\frac{x}{|x|}\right) - \ln |x|.$$

(c) If Σ is star-shaped, then

$$G_\Sigma(\omega) = \frac{1}{2} [\ln(\ell_\Sigma(\omega)) + \ln(\ell_\Sigma(-\omega))] \quad (2.4)$$

for each $\omega \in \mathbb{S}^{d-1}$.

Proof. Let $x \in \mathbb{R}^d \setminus \{0\}$. Then point (a) follows from the equalities

$$\begin{aligned} \lim_{\varepsilon \searrow 0} \left(\int_{\varepsilon}^{+\infty} \frac{d\mu}{\mu} \mathbb{1}_{\Sigma}(\mu x) + \ln \varepsilon \right) &= \int_1^{+\infty} \frac{d\mu}{\mu} \mathbb{1}_{\Sigma}(\mu x) + \lim_{\varepsilon \searrow 0} \int_{\varepsilon}^1 \frac{d\mu}{\mu} [\mathbb{1}_{\Sigma}(\mu x) - 1] \\ &= \int_1^{+\infty} \frac{d\mu}{\mu} \mathbb{1}_{\Sigma}(\mu x) + \int_0^1 \frac{d\mu}{\mu} [\mathbb{1}_{\Sigma}(\mu x) - 1]. \end{aligned}$$

Furthermore we have for $\lambda > 0$

$$\begin{aligned} R_{\Sigma}(\lambda x) &= \lim_{\varepsilon \searrow 0} \left(\int_{\varepsilon}^{+\infty} \frac{d\mu}{\mu} \mathbb{1}_{\Sigma}(\mu \lambda x) + \ln \varepsilon \right) \\ &= \lim_{\varepsilon \searrow 0} \left(\int_{\lambda \varepsilon}^{+\infty} \frac{d\mu}{\mu} \mathbb{1}_{\Sigma}(\mu x) + \ln(\lambda \varepsilon) - \ln \lambda \right) \\ &= R_{\Sigma}(x) - \ln \lambda, \end{aligned}$$

which proves point (b). Finally point (c) follows from a direct computation. \square

We give now some properties of the functions R_{Σ} and G_{Σ} , which follow easily from Lemma 2.2.

Remark 2.3. (a) Let us consider $\mathbb{R}_+^* :=]0, +\infty[$ endowed with the multiplication as a Lie group with Haar measure $\frac{d\mu}{\mu}$. Then R_{Σ} is the (renormalized) average of $\mathbb{1}_{\Sigma}$ with respect to the action of \mathbb{R}_+^* on \mathbb{R}^d .

(b) If Σ is equal to the unit open ball $\mathcal{B} := \{x \in \mathbb{R}^d \mid |x| < 1\}$, then we have

$$G_{\mathcal{B}}(x) = -\ln |x|, \quad x \in \mathbb{R}^d \setminus \{0\}. \quad (2.5)$$

(c) To each set Σ one can associate a unique symmetric star-shaped set $\tilde{\Sigma}$ such that

$$G_{\Sigma} = G_{\tilde{\Sigma}}.$$

Indeed it suffices to take the symmetric star-shaped set $\tilde{\Sigma}$ defined by the even, strictly positive, continuous function ℓ on \mathbb{S}^{d-1} given by $\ell(\omega) := \exp(G_{\Sigma}(\omega))$.

3 Symmetrized time delay in classical scattering

In this section we study the symmetrized time delay in classical scattering defined in terms of sojourn times in the sets Σ_r . For simplicity we restrict ourselves to scattering by compactly supported potentials. Under a natural assumption on the potential we show that Assumption 2.1 is a necessary condition for the existence of the symmetrized time delay. We also prove the existence of the symmetrized time delay for symmetric star-shaped regions with C^2 -boundary and show that the symmetrized time delay and the usual time delay are equal if the sets Σ_r are spherical.

Let Σ be a bounded open set containing 0 and let V be a real C^2 -potential with compact support. A *scattering trajectory* for V is a map $\Phi : \mathbb{R} \ni t \mapsto (x(t), p(t)) \in \mathbb{R}^{2d}$ solution of $\dot{x}(t) = p(t)$, $\dot{p}(t) = -\nabla V(x(t))$ satisfying $|x(t)| \rightarrow \infty$ as $t \rightarrow \pm\infty$ and $E := \frac{1}{2}p^2(t) + V(x(t)) > 0$. Since V has compact support there exist times $t_- \leq t_+$, positions x_\pm , and momenta p_\pm with $|p_\pm| = \sqrt{2E}$ such that

$$x(t) = x_\pm + (t - t_\pm)p_\pm \quad \text{if } \pm t \geq \pm t_\pm.$$

Thus the *incoming* and *outgoing free trajectories* Φ_\pm associated to the scattering trajectory Φ are given by

$$x_\pm(t) := x_\pm + (t - t_\pm)p_\pm, \quad p_\pm(t) := p_\pm.$$

For each $r \in \mathbb{R} \setminus \{0\}$ we define the characteristic function

$$\chi_r(x) := \mathbb{1}_\Sigma(r^{-1}x), \quad x \in \mathbb{R}^d,$$

and for each suitable trajectory $\Psi : \mathbb{R} \ni t \mapsto (y(t), v(t)) \in \mathbb{R}^{2d}$ we set

$$T_r(\Psi) := \int_{-\infty}^{+\infty} dt \chi_r(y(t)).$$

Then the sojourn time in Σ_r of the scattering trajectory Φ is $T_r(\Phi)$ and the sojourn times in Σ_r of the incoming and outgoing free trajectories Φ_\pm are $T_r(\Phi_\pm)$. The usual time delay of the scattering process for Σ_r is defined as

$$\tau_r^{\text{in}}(\Phi) := T_r(\Phi) - T_r(\Phi_-).$$

It is known [12, Sec. 2] that $\lim_{r \rightarrow \infty} \tau_r^{\text{in}}(\Phi)$ exists only if Σ is spherical. In this case $\tau_r^{\text{in}}(\Phi)$ converges to the classical analogue $\tau^{\text{cl}}(\Phi)$ of the Eisenbud-Wigner time delay [11].

One can also define the *symmetrized time delay*:

$$\tau_r(\Phi) := T_r(\Phi) - \frac{1}{2} [T_r(\Phi_+) + T_r(\Phi_-)].$$

Remark 3.1. Consider the mapping f (of full time reversal) $f : \Phi \mapsto \Phi^{\text{rev}}$, where $\Phi^{\text{rev}} : t \mapsto (x(-t), -p(-t))$, and set

$$\tau_r^{\text{out}}(\Phi) := (\tau_r^{\text{in}} \circ f)(\Phi).$$

Since $T_r(\Phi^{\text{rev}}) = T_r(\Phi)$ and $T_r((\Phi^{\text{rev}})_-) = T_r(\Phi_+)$, we have

$$\tau_r(\Phi) = \frac{1}{2} [\tau_r^{\text{in}}(\Phi) + \tau_r^{\text{out}}(\Phi)].$$

Thus τ_r is the mean value of the usual time delay τ_r^{in} and of the time delay τ_r^{out} corresponding to the time reversed scattering process. In particular τ_r is invariant under full time reversal, namely one has $\tau_r \circ f = \tau_r$ since f is an involution.

The following assumption concerns the potential V . Given $\omega \in \mathbb{S}^{d-1}$, $S(\omega) \subset \mathbb{S}^{d-1}$ stands for the set of all outgoing directions of scattering trajectories with incoming direction ω .

Assumption 3.2. *For all $\omega \in \mathbb{S}^{d-1}$, $S(\omega)$ contains a neighborhood of ω .*

We believe that Assumption 3.2 should be satisfied by typical nonzero potentials. A possible way to prove it would be to show that at high energies the differential of the outgoing direction with respect to the incoming impact parameter is non zero.

Theorem 3.3. (a) *Assume that V satisfies Assumption 3.2. Then $\lim_{r \rightarrow \infty} \tau_r$ exists and is finite for all scattering trajectories only if Σ satisfies Assumption 2.1.*

(b) *If Σ is symmetric and star-shaped with a C^2 -boundary, then $\lim_{r \rightarrow \infty} \tau_r$ exists and is finite for all scattering trajectories.*

(c) *If Σ is spherical, then $\lim_{r \rightarrow \infty} [\tau_r(\Phi) - \tau_r^{\text{in}}(\Phi)] = 0$.*

Proof. (a) A direct calculation gives

$$T_r(\Phi) = \frac{1}{\sqrt{2E}} \int_{-\infty}^0 dt \chi_r(x_- + t\omega_-) + \int_{t_-}^{t_+} dt \chi_r(x(t)) + \frac{1}{\sqrt{2E}} \int_0^{+\infty} dt \chi_r(x_+ + t\omega_+)$$

and

$$T_r(\Phi_{\pm}) = \frac{1}{\sqrt{2E}} \int_{-\infty}^{+\infty} dt \chi_r(x_{\pm} + t\omega_{\pm}),$$

where $\omega_{\pm} := p_{\pm}/|p_{\pm}|$, $|p_{\pm}| = \sqrt{2E}$. It follows that

$$\tau_r(\Phi) = \int_{t_-}^{t_+} dt \chi_r(x(t)) + \frac{1}{\sqrt{8E}} [I(x_+, \omega_+, r) - I(x_-, \omega_-, r)], \quad (3.6)$$

where

$$I(x, \omega, r) := \int_{-\infty}^{+\infty} dt \operatorname{sgn}(t) \chi_r(x + t\omega), \quad x \in \mathbb{R}^d, \omega \in \mathbb{S}^{d-1}.$$

The first term in the r.h.s. of Equation (3.6) tends to $t_+ - t_-$, so it is sufficient to consider the convergence of the second term. Since $I(x, \omega, r) = rI(r^{-1}x, \omega, 1)$ and $\lim_{x \rightarrow 0} I(x, \omega, 1) = I(0, \omega, 1)$, we get

$$I(x, \omega, r) = rI(0, \omega, 1) + o(r),$$

and

$$I(x_+, \omega_+, r) - I(x_-, \omega_-, r) = r [I(0, \omega_+, 1) - I(0, \omega_-, 1)] + o(r).$$

Thus $\lim_{r \rightarrow \infty} \tau_r(\Phi)$ exists only if

$$I(0, \omega_+, 1) = I(0, \omega_-, 1).$$

Varying the scattering trajectory keeping ω_- fixed, we deduce from Assumption 3.2 that $I(0, \cdot, 1)$ is locally constant on \mathbb{S}^{d-1} , and thus constant by connexity. Since $I(0, \cdot, 1)$ is

an odd function we get the equality $I(0, \cdot, 1) = 0$, which is equivalent to Assumption 2.1.

(b) Suppose that Σ is symmetric star-shaped with a C^2 -boundary, and let $\ell_\Sigma : \mathbb{S}^{d-1} \rightarrow (0, \infty)$ be the function associated to Σ as in Section 2. We extend ℓ_Σ to \mathbb{R}^d by homogeneity. Given $x \in \mathbb{R}^d$ orthogonal to $\omega \in \mathbb{S}^{d-1}$, we denote by $s_\pm \equiv s_\pm(r^{-1}x, \omega)$ the entrance and exit times of the free trajectory $t \mapsto (r^{-1}x + t\omega, \omega)$ in Σ (s_\pm exist if r is large enough). Since $\ell_\Sigma(\omega) = \ell_\Sigma(-\omega)$, the times s_\pm satisfy $s_\pm(0, \omega) = \pm\ell_\Sigma(\omega)$ and are solutions of the equations

$$s_\pm = \pm (\ell_\Sigma(r^{-1}x + s_\pm\omega)^2 - r^{-2}x^2)^{1/2}.$$

Doing a Taylor expansion around $r^{-1} = 0$, we get

$$s_\pm(r^{-1}x, \omega) = \pm\ell_\Sigma(\omega) \pm r^{-1}\ell_\Sigma(\omega)^{-1}x \cdot (\nabla\ell_\Sigma)(\pm\omega) + O(r^{-2}). \quad (3.7)$$

Thus, for each $x \in \mathbb{R}^d$, $\omega \in \mathbb{S}^{d-1}$, and $x^\perp := x - (x \cdot \omega)\omega$, we get

$$I(r^{-1}x^\perp, \omega, 1) = s_+(r^{-1}x^\perp, \omega) + s_-(r^{-1}x^\perp, \omega) = O(r^{-1}).$$

This implies that

$$\begin{aligned} I(x, \omega, r) &= rI(r^{-1}x, \omega, 1) = r \int_{-\infty}^{+\infty} dt \operatorname{sgn}(t - r^{-1}(x \cdot \omega)) \mathbb{1}_\Sigma(r^{-1}x^\perp + t\omega) \\ &= rI(r^{-1}x^\perp, \omega, 1) + O(1) \\ &= O(1), \end{aligned}$$

and the existence of $\lim_{r \rightarrow \infty} \tau_r(\Phi)$ follows from Equation (3.6).

(c) If Σ is spherical, then Equation (3.7) gives

$$T_r(\Phi_\pm) = \frac{r}{\sqrt{2E}} + O(r^{-1}).$$

This implies that $\lim_{r \rightarrow \infty} [\tau_r(\Phi) - \tau_r^{\text{in}}(\Phi)] = 0$. □

4 Symmetrized time delay in quantum scattering

4.1 Sojourn times

In this section we gather some properties of the (quantum) sojourn times associated to the free Hamiltonian $H_0 = -\frac{1}{2}\Delta$ and the full Hamiltonian H in $\mathcal{H} = L^2(\mathbb{R}^d)$. We first recall some definitions.

Σ is a bounded open set in \mathbb{R}^d containing 0, and $\Sigma_r = \{rx \mid x \in \Sigma\}$. We write $\mathbb{1}_{H_0}(\cdot)$ for the spectral measure of H_0 and Q for the (vector) position operator in \mathcal{H} . We set $\langle \cdot \rangle := \sqrt{1 + |\cdot|^2}$, and we always assume that:

Assumption 4.1. *The wave operators W^\pm exist and are complete. The projections $\chi_r(Q)$ are locally H -smooth on $]0, +\infty[\setminus\sigma_{\text{pp}}(H)$.*

For latter use we introduce the following definition:

Definition 4.2. *Let $s \geq 0$, then*

$$\mathcal{D}_s := \{\varphi \in \mathcal{D}(\langle Q \rangle^s) \mid \mathbb{1}_{H_0}(J)\varphi = \varphi \text{ for some compact set } J \text{ in }]0, +\infty[\setminus\sigma_{\text{pp}}(H)\}.$$

It is clear that \mathcal{D}_s is dense in \mathcal{H} and that $\mathcal{D}_{s_1} \subset \mathcal{D}_{s_2}$ if $s_1 \geq s_2$.

For $r > 0$ and an appropriate scattering state $\varphi \in \mathcal{H}$, we define the *free sojourn time*

$$T_r^0(\varphi) := \int_{-\infty}^{+\infty} dt \|\chi_r(Q) e^{-itH_0} \varphi\|^2$$

and the *full sojourn time*

$$T_r(\varphi) := \int_{-\infty}^{+\infty} dt \|\chi_r(Q) e^{-itH} W^- \varphi\|^2.$$

Due to Assumption 4.1, one shows easily that these times are finite if $\varphi \in \mathcal{D}_0$. The *time delay* of the scattering process with incoming state $\varphi \in \mathcal{D}_0$ for Σ_r is then defined as

$$\tau_r^{\text{in}}(\varphi) := T_r(\varphi) - T_r^0(\varphi).$$

Since $S\mathcal{D}_0 \subset \mathcal{D}_0$ one can also define the *symmetrized time delay* of the scattering process with incoming state $\varphi \in \mathcal{D}_0$:

$$\tau_r(\varphi) := T_r(\varphi) - \frac{1}{2} [T_r^0(\varphi) + T_r^0(S\varphi)].$$

Finally we define for each $r > 0$ the auxiliary sojourn time $\tau_r^{\text{free}}(\varphi)$ (see [14, Sec. 2.1])

$$\begin{aligned} \tau_r^{\text{free}}(\varphi) := & \frac{1}{2} \int_{-\infty}^0 dt (\|\chi_r(Q) e^{-itH_0} \varphi\|^2 - \|\chi_r(Q) e^{-itH_0} S\varphi\|^2) \\ & + \frac{1}{2} \int_0^{+\infty} dt (\|\chi_r(Q) e^{-itH_0} S\varphi\|^2 - \|\chi_r(Q) e^{-itH_0} \varphi\|^2), \end{aligned} \quad (4.8)$$

which is also finite if $\varphi \in \mathcal{D}_0$.

Lemma 4.3. *Suppose that Assumption 4.1 holds and let $\varphi \in \mathcal{D}_0$ be such that*

$$\|(W^- - \mathbb{1}) e^{-itH_0} \varphi\| \in L^1(\mathbb{R}_-, dt) \quad (4.9)$$

and

$$\|(W^+ - \mathbb{1}) e^{-itH_0} S\varphi\| \in L^1(\mathbb{R}_+, dt). \quad (4.10)$$

Then

$$\lim_{r \rightarrow \infty} [\tau_r(\varphi) - \tau_r^{\text{free}}(\varphi)] = 0.$$

Proof. For $t \in \mathbb{R}$, set

$$f_-(t) := \|e^{-itH} W^- \varphi - e^{-itH_0} \varphi\| \quad \text{and} \quad f_+(t) := \|e^{-itH} W^+ \varphi - e^{-itH_0} S\varphi\|.$$

We know from Hypotheses (4.9) and (4.10) that $f_{\pm} \in L^1(\mathbb{R}_{\pm})$. Using the inequality

$$|\|u\|^2 - \|v\|^2| \leq \|u - v\|(\|u\| + \|v\|), \quad u, v \in \mathcal{H}, \quad (4.11)$$

we obtain the estimates

$$\begin{aligned} |\|\chi_r(Q) e^{-itH} W^- \varphi\|^2 - \|\chi_r(Q) e^{-itH_0} \varphi\|^2| &\leq 2f_-(t)\|\varphi\|, \\ |\|\chi_r(Q) e^{-itH} W^+ \varphi\|^2 - \|\chi_r(Q) e^{-itH_0} S\varphi\|^2| &\leq 2f_+(t)\|\varphi\|. \end{aligned}$$

Since $s\text{-}\lim_{r \rightarrow +\infty} \chi_r(Q) = \mathbb{1}$, then the scalars on the l.h.s. above converge to 0 as $r \rightarrow +\infty$. Thus the claim follows from (4.8) and Lebesgue's dominated convergence theorem. \square

4.2 Time reversal

We now collect some elementary remarks related to time reversal for the (complete) scattering system $\{H_0, H\}$.

Time reversal is implemented by the antiunitary involution

$$\mathcal{H} \ni \varphi \mapsto \overline{\varphi}.$$

The Hamiltonian H is invariant under time reversal if

$$H\overline{\varphi} = \overline{H\varphi}, \quad \varphi \in \mathcal{D}(H).$$

In such a case one has the identities $e^{-itH} \overline{\varphi} = \overline{e^{itH} \varphi}$, $W^{\pm} \overline{\varphi} = \overline{W^{\mp} \varphi}$ and

$$S\overline{\varphi} = \overline{S^{-1}\varphi} \quad (4.12)$$

for each $\varphi \in \mathcal{H}$. Consider the bijection

$$f: \mathcal{H} \rightarrow \mathcal{H}, \quad \varphi \mapsto \overline{S\varphi},$$

which we call *full time reversal*. The map f corresponds to time reversal for the *full scattering process*, i.e. it interchanges past and future scattering data and reverses the direction of time. Furthermore one sees easily from (4.12) that f is an antiunitary involution.

In order to give a rigorous interpretation of full time reversal we introduce the space \mathcal{E} of *scattering trajectories*, i.e. the space of continuous maps

$$\mathbb{R} \ni t \mapsto \Phi(t) \in \mathcal{H},$$

such that

$$i(\partial_t \Phi)(t) = H\Phi(t) \quad \forall t \in \mathbb{R} \quad (\text{in the weak sense}) \quad \text{and} \quad w\text{-}\lim_{t \rightarrow \pm\infty} \Phi(t) = 0.$$

The space \mathcal{E} is invariant under the involution

$$R : \mathcal{E} \rightarrow \mathcal{E}, \quad (R\Phi)(t) := \overline{\Phi(-t)}.$$

One can associate to a trajectory $\Phi \in \mathcal{E}$ a vector $\varphi := T(\Phi) \in \mathcal{H}$ defined by the constraint

$$s\text{-}\lim_{t \rightarrow -\infty} (\Phi(t) - e^{-itH_0} \varphi) = 0.$$

Due to the completeness of the wave operators we know that $T : \mathcal{E} \rightarrow \mathcal{H}$ is bijective, and we have

$$f(\varphi) = (TRT^{-1})(\varphi), \quad \varphi \in \mathcal{H}. \quad (4.13)$$

Equation (4.13) provides a rigorous meaning to full time reversal as a map interchanging past and future scattering data and reversing the direction of time.

Lemma 4.4. *Assume that H is invariant under time reversal, and set*

$$\tau_r^{\text{out}}(\varphi) := (\tau_r^{\text{in}} \circ f)(\varphi).$$

Then one has the equalities

$$\tau_r(\varphi) = \frac{1}{2} [\tau_r^{\text{in}}(\varphi) + \tau_r^{\text{out}}(\varphi)] \quad \text{and} \quad \tau_r(\varphi) = (\tau_r \circ f)(\varphi). \quad (4.14)$$

Thus $\tau_r(\varphi)$ is the mean value of the usual time delay $\tau_r^{\text{in}}(\varphi)$ and of the time delay $\tau_r^{\text{out}}(\varphi)$ corresponding to the time reversed scattering process. In particular $\tau_r(\varphi)$ is invariant under full time reversal.

Proof. Since H_0 is invariant under time reversal, one gets

$$T_r^0(\varphi) = T_r^0(\overline{\varphi}).$$

This together with time reversal invariance of H yields

$$T_r(\overline{S\varphi}) = T_r(\varphi).$$

Thus

$$\tau_r^{\text{out}}(\varphi) = \tau_r^{\text{in}}(\overline{S\varphi}) = T_r(\overline{S\varphi}) + T_r^0(\overline{S\varphi}) = T_r(\varphi) + T_r^0(S\varphi),$$

which implies the first identity in (4.14). The second identity follows from the fact that f is an involution. \square

4.3 Time delay

In the present section we shall give the proof of the existence of the symmetrized time delay. We first fix some notation. If A, B are two symmetric operators, then we set for each $\varphi \in \mathcal{D}(A) \cap \mathcal{D}(B)$:

$$(\varphi, [A, B]\varphi) := (A\varphi, B\varphi) - (B\varphi, A\varphi).$$

If q is a quadratic form with domain $\mathcal{D}(q)$, and S is unitary, then we set for each $\varphi \in \mathcal{D}(q) \cap S^{-1}\mathcal{D}(q)$:

$$(\varphi, S^*[q, S]\varphi) := q(S\varphi) - q(\varphi).$$

If A is an operator with domain $\mathcal{D}(A)$ and S is unitary, then we define the operator $S^*[A, S]$ with domain $\mathcal{D}(A) \cap S^{-1}\mathcal{D}(A)$ by

$$S^*[A, S] := S^*AS - A.$$

We also recall that the function G_Σ was introduced in Section 2 and that $\mathcal{D}_2 \subset \mathcal{D}(Q^2) \cap \mathcal{D}(G_\Sigma(P))$. Therefore the quadratic form $i[Q^2, G_\Sigma(P)]$ is well defined on \mathcal{D}_2 .

The proof of the next proposition can be found in the appendix.

Proposition 4.5. *Let Σ satisfy Assumption 2.1. Suppose that Assumption 4.1 is verified. Then we have for all $\varphi \in \mathcal{D}_2$ the equality*

$$\begin{aligned} & \lim_{r \rightarrow \infty} \int_0^{+\infty} dt (\varphi, (e^{itP^2/2} \chi_r(Q) e^{-itP^2/2} - e^{-itP^2/2} \chi_r(Q) e^{itP^2/2}) \varphi) \\ &= -(\varphi, i[Q^2, G_\Sigma(P)]\varphi). \end{aligned} \quad (4.15)$$

We are now in a position to give the proof of our main theorem. It involves the operator

$$A_0 := \frac{1}{2} \left(\frac{P}{|P|^2} \cdot Q + Q \cdot \frac{P}{|P|^2} \right),$$

which is well-defined and symmetric on \mathcal{D}_1 .

Theorem 4.6. *Let Σ satisfy Assumption 2.1. Suppose that Assumption 4.1 is verified. Let $\varphi \in \mathcal{D}_2$ satisfy (4.9), (4.10) and $S\varphi \in \mathcal{D}_2$. Then the limit $\tau_\Sigma(\varphi) = \lim_{r \rightarrow \infty} \tau_r(\varphi)$ exists, and one has*

$$\tau_\Sigma(\varphi) = -\frac{1}{2} (\varphi, S^*[i[Q^2, G_\Sigma(\frac{P}{|P|})], S]\varphi) - (\varphi, S^*[A_0, S]\varphi). \quad (4.16)$$

The quadratic form $i[Q^2, G_\Sigma(\frac{P}{|P|})]$ and the operator A_0 are well defined on \mathcal{D}_2 , so all the commutators in the above formula are well defined since $\varphi, S\varphi \in \mathcal{D}_2$.

Proof. The expression (4.8) for $\tau_r^{\text{free}}(\varphi)$ can be rewritten as

$$\begin{aligned} \tau_r^{\text{free}}(\varphi) &= -\frac{1}{2} \int_0^{+\infty} dt (\varphi, (e^{itP^2/2} \chi_r(Q) e^{-itP^2/2} - e^{-itP^2/2} \chi_r(Q) e^{itP^2/2}) \varphi) \\ &\quad + \frac{1}{2} \int_0^{+\infty} dt (S\varphi, (e^{itP^2/2} \chi_r(Q) e^{-itP^2/2} - e^{-itP^2/2} \chi_r(Q) e^{itP^2/2}) S\varphi). \end{aligned}$$

Applying Proposition 4.5, we get

$$\begin{aligned} \lim_{r \rightarrow +\infty} \tau_r^{\text{free}}(\varphi) &= \frac{1}{2} (\varphi, i[Q^2, G_\Sigma(P)]\varphi) - \frac{1}{2} (S\varphi, i[Q^2, G_\Sigma(P)]S\varphi) \\ &= -\frac{1}{2} (\varphi, S^*[i[Q^2, G_\Sigma(P)], S]\varphi). \end{aligned}$$

By Lemma 2.2.(b), we have $G_\Sigma(P) = G_\Sigma\left(\frac{P}{|P|}\right) - \ln |P|$, and we know from [2, Sec. 2] that

$$\frac{i}{2}[Q^2, -\ln |P|] = A_0,$$

as quadratic forms on \mathcal{D}_2 . This yields

$$\lim_{r \rightarrow +\infty} \tau_r^{\text{free}}(\varphi) = -\frac{1}{2}(\varphi, S^* [i[Q^2, G_\Sigma\left(\frac{P}{|P|}\right)], S] \varphi) - (\varphi, S^* [A_0, S] \varphi).$$

We conclude by using Lemma 4.3. \square

Remark 4.7. *The second term in Formula (4.16) coincides with the usual value of time delay; it is equal to the limit (for Σ spherical) of $\tau_r^{\text{in}}(\varphi)$ as $r \rightarrow +\infty$ (see [2, Prop. 1]). The first term is a new contribution to time delay determined by the shape of the set Σ . If Σ is spherical, this contribution vanishes due to Remark 2.3.(b), and then one gets (under the hypotheses of Theorem 4.6) the equality*

$$\lim_{r \rightarrow \infty} \tau_r(\varphi) = \lim_{r \rightarrow \infty} \tau_r^{\text{in}}(\varphi).$$

Remark 4.8. *Under the hypotheses of Theorem 4.6, the two following facts are true whenever Σ is an open bounded set containing the origin (see [12, Sec. 3]):*

(a) *The equality $\lim_{r \rightarrow \infty} [\tau_r(\varphi) - \tau_r^{\text{free}}(\varphi)] = 0$ holds.*

(b) *The difference*

$$\begin{aligned} & \int_0^{+\infty} dt (S\varphi, [e^{itP^2} \chi_r(Q) e^{-itP^2} - e^{-itP^2} \chi_r(Q) e^{itP^2}], S] \varphi) \\ & - r \int_0^{+\infty} du (S\varphi, [|H_0|^{-1/2} (\mathbb{1}_\Sigma(u\frac{P}{|P|}) - \mathbb{1}_\Sigma(-u\frac{P}{|P|}))], S] \varphi) \end{aligned}$$

remains bounded as $r \rightarrow \infty$.

The integrand in the second term in (b) can be written as

$$(S\varphi, |H_0|^{-1/2} [M(P), S] \varphi),$$

where

$$M(x) = |x| \int_0^{+\infty} d\mu [\mathbb{1}_\Sigma(\mu x) - \mathbb{1}_\Sigma(-\mu x)], \quad x \in \mathbb{R}^d.$$

The combination of facts (a) and (b) shows that $\tau_r(\varphi)$ can have a limit for φ in a dense set $\mathcal{E} \subset \mathcal{H}$ only if

$$(S\varphi, |H_0|^{-1/2} [M(P), S] \varphi) = 0, \quad \forall \varphi \in \mathcal{E},$$

which implies that $[M(P), S] = 0$. Therefore, if the scattering operator S has no other symmetry than $[S, P^2] = 0$, one has $M(P) = F(P^2)$ for some function F , and it follows that $M \equiv 0$ since $M(x) = -M(-x)$. In consequence $\tau_r(\varphi)$ can have a limit for φ in a dense set \mathcal{E} only if Σ satisfies Assumption 2.1.

Remark 4.9. One could also consider time delay for sets Σ_r translated by a vector $a \in \mathbb{R}^d$. Obviously this is equivalent to determining the time delay (4.16) when the origin of the spatial coordinate system is translated to the point a . In this case one has

$$\begin{aligned}\varphi &\mapsto \varphi_a := e^{iP \cdot a} \varphi, \\ S &\mapsto S_a := e^{iP \cdot a} S e^{-iP \cdot a},\end{aligned}$$

and $\tau_\Sigma(\varphi)$ becomes

$$\tau_\Sigma^a(\varphi) := -\frac{1}{2}(\varphi_a, S_a^* [i[Q^2, G_\Sigma(\frac{P}{|P|})], S_a] \varphi_a) - (\varphi_a, S_a^* [A_0, S_a] \varphi_a).$$

Using the formulas

$$\begin{aligned}e^{-iP \cdot a} [Q^2, G_\Sigma(\frac{P}{|P|})] e^{iP \cdot a} &= [Q^2, G_\Sigma(\frac{P}{|P|})] - 2a \cdot [Q, G_\Sigma(\frac{P}{|P|})], \\ e^{-iP \cdot a} A_0 e^{iP \cdot a} &= A_0 - a \cdot \frac{P}{|P|^2},\end{aligned}$$

one gets

$$\tau_\Sigma^a(\varphi) = \tau_\Sigma(\varphi) + a \cdot (\varphi, S^* [i[Q, G_\Sigma(\frac{P}{|P|})], S] \varphi) + a \cdot (\varphi, S^* [\frac{P}{|P|^2}, S] \varphi). \quad (4.17)$$

Due to its very definition time delay given by Formula (4.17) is clearly covariant under spatial translations.

4.4 Stationary formulas

In the sequel we derive stationary formulas for the symmetrized time delay in the case of scattering by a short-ranged potential. We refer to [5] for the treatment of this issue in the case of the usual time delay.

Suppose that Σ satisfy Assumption 2.1. Then we know from Remark 2.3.(c) that there exists a symmetric star-shaped set $\tilde{\Sigma}$ such that

$$\tau_\Sigma(\varphi) = \tau_{\tilde{\Sigma}}(\varphi),$$

for φ satisfying the hypotheses of Theorem 4.6. Thus with no loss of generality we may assume that Σ is symmetric and star-shaped. We also assume that the boundary $\partial\Sigma$ of Σ is a C^2 hypersurface, so that the functions ℓ_Σ and G_Σ (see Formulas (2.3) and (2.4)) associated to Σ are C^2 . In such a case one has

$$\frac{i}{2} [Q^2, G_\Sigma(\frac{P}{|P|})] = -\frac{1}{2} [Q \cdot \nabla G_\Sigma(\frac{P}{|P|}) + \nabla G_\Sigma(\frac{P}{|P|}) \cdot Q] =: B_\Sigma,$$

as quadratic forms on \mathcal{D}_2 (note that B_Σ is a well-defined symmetric operator on \mathcal{D}_1). Thus we can rewrite $\tau_\Sigma(\varphi)$ as

$$\tau_\Sigma(\varphi) = -(\varphi, S^* [B_\Sigma, S] \varphi) - (\varphi, S^* [A_0, S] \varphi). \quad (4.18)$$

Let $\mathcal{U} : \mathbb{L}^2(\mathbb{R}^d) \rightarrow \int_{\mathbb{R}_+}^{\oplus} d\lambda \mathbb{L}^2(\mathbb{S}^{d-1})$ be the spectral transformation for H_0 , *i.e.* the unitary mapping defined by

$$(\mathcal{U}\varphi)(\lambda, \omega) = (2\lambda)^{(d-2)/4} (\mathcal{F}\varphi)(\sqrt{2\lambda}\omega),$$

where \mathcal{F} denotes the Fourier transform. One has

$$\mathcal{U}H_0\mathcal{U}^{-1} = \int_{\mathbb{R}_+}^{\oplus} d\lambda \lambda \quad \text{and} \quad \mathcal{U}S\mathcal{U}^{-1} = \int_{\mathbb{R}_+}^{\oplus} d\lambda S(\lambda),$$

where $\{S(\lambda)\}_{\lambda \geq 0} \subset \mathcal{B}(\mathbb{L}^2(\mathbb{S}^{d-1}))$ is the scattering matrix for the pair $\{H_0, H\}$. For shortness we shall set $\varphi(\lambda) := (T\varphi)(\lambda, \cdot) \in \mathbb{L}^2(\mathbb{S}^{d-1})$.

If the interaction $V := H - H_0$ is a potential that decays faster than $|x|^{-2}$ at infinity, then there exists a dense set $\mathcal{E} \subset \mathcal{H}$ such that the hypotheses of Theorem 4.6 are satisfied for any $\varphi \in \mathcal{E}$ (a precise definition of V and \mathcal{E} can be found in [3, Prop. 3], see also [7, 8] for related results). Furthermore the function $\lambda \mapsto S(\lambda)$ is strongly continuously differentiable on \mathcal{E} , and the second term in (4.18) is equal to the Eisenbud-Wigner time delay for any $\varphi \in \mathcal{E}$:

$$-(\varphi, S^*[A_0, S]\varphi) = -i \int_0^\infty d\lambda \left(\varphi(\lambda), S(\lambda)^* \left(\frac{dS(\lambda)}{d\lambda} \right) \varphi(\lambda) \right)_{\mathbb{L}^2(\mathbb{S}^{d-1})} \equiv (\varphi, \tau_{E-W}\varphi).$$

Let us now consider the first term in (4.18). Since the function $x \mapsto G_\Sigma\left(\frac{x}{|x|}\right)$ is homogeneous of degree 0, one has

$$x \cdot (\nabla G_\Sigma)\left(\frac{x}{|x|}\right) = 0,$$

namely the vector field $(\nabla G_\Sigma)\left(\frac{x}{|x|}\right)$ is orthogonal to the radial direction. In fact a direct calculation shows that

$$\mathcal{U}B_\Sigma\mathcal{U}^{-1} = \int_{\mathbb{R}_+}^{\oplus} d\lambda \lambda^{-1} b_\Sigma(\omega, \partial_\omega),$$

where $b_\Sigma(\omega, \partial_\omega)$ is a symmetric first order differential operator on \mathbb{S}^{d-1} with C^1 coefficients. Therefore the operator $\mathcal{U}B_\Sigma\mathcal{U}^{-1}$ is essentially selfadjoint on $\mathcal{U}\mathcal{D}_1$, and its closure is decomposable in the spectral representation of H_0 , *i.e.*

$$\overline{\mathcal{U}B_\Sigma\mathcal{U}^{-1}} = \mathcal{U}\overline{B_\Sigma}\mathcal{U}^{-1} \equiv \int_{\mathbb{R}_+}^{\oplus} d\lambda B_\Sigma(\lambda).$$

This yields the equality

$$\begin{aligned} \tau_\Sigma(\varphi) &= - \int_0^{+\infty} d\lambda (\varphi(\lambda), S^*(\lambda)[B_\Sigma(\lambda), S(\lambda)]\varphi(\lambda))_{\mathbb{L}^2(\mathbb{S}^{d-1})} \\ &\quad - i \int_0^{+\infty} d\lambda \left(\varphi(\lambda), S(\lambda)^* \left(\frac{dS(\lambda)}{d\lambda} \right) \varphi(\lambda) \right)_{\mathbb{L}^2(\mathbb{S}^{d-1})}. \end{aligned}$$

In consequence the time delay (4.18) is the sum of two contributions, each of these being the expectation value of an operator decomposable in the spectral representation of H_0 .

Appendix

Proof of Proposition 4.5. (i) For any $F \in L^\infty(\mathbb{R}^d)$ and $s \in \mathbb{R}$ one has

$$\begin{aligned} e^{isP^2/2} F(Q) e^{-isP^2/2} &= F(Q + sP), \\ e^{-isQ^2/2} F(P) e^{isQ^2/2} &= F(P + sQ), \end{aligned} \quad (4.19)$$

which imply the identity

$$e^{itP^2/2} F(Q) e^{-itP^2/2} = Z_{-1/t} F(tP) Z_{1/t}, \quad (4.20)$$

where $t \in \mathbb{R}^*$ and $Z_\tau := e^{i\tau Q^2/2}$. Formula (4.20) and the change of variables $\mu = rt^{-1}$, $\nu = r^{-1}$, lead to the equalities

$$\begin{aligned} &\int_0^{+\infty} dt (\varphi, (e^{itP^2/2} \chi_r(Q) e^{-itP^2/2} - e^{-itP^2/2} \chi_r(Q) e^{itP^2/2}) \varphi) \\ &= \int_0^{+\infty} \frac{d\mu}{\nu\mu^2} (\varphi, (Z_{-\nu\mu} \chi_\mu(P) Z_{\nu\mu} - Z_{\nu\mu} \chi_{-\mu}(P) Z_{-\nu\mu}) \varphi). \end{aligned}$$

One has also

$$\int_0^{+\infty} \frac{d\mu}{\mu^2} [\chi_\mu(P) - \chi_{-\mu}(P)] = \int_0^{+\infty} ds [\chi(sP) - \chi(-sP)] = 0$$

due to Assumption 2.1. Hence the l.h.s of (4.15) can be written as

$$K_\infty(\varphi) := \lim_{\nu \searrow 0} \int_0^{+\infty} d\mu K_{\nu,\mu}(\varphi), \quad (4.21)$$

where

$$\begin{aligned} K_{\nu,\mu}(\varphi) &:= \frac{1}{\nu\mu^2} (\varphi, [Z_{-\nu\mu} \chi_\mu(P) Z_{\nu\mu} - \chi_\mu(P)] \varphi) \\ &\quad - \frac{1}{\nu\mu^2} (\varphi, [Z_{\nu\mu} \chi_{-\mu}(P) Z_{-\nu\mu} - \chi_{-\mu}(P)] \varphi). \end{aligned}$$

(ii) To prove the statement, we shall show that one may interchange the limit and the integral in (4.21), by invoking Lebesgue's dominated convergence theorem. This will be done in (iii) below. If one assumes that this interchange is justified for the moment, then direct calculations give

$$\begin{aligned} K_\infty(\varphi) &= \int_0^{+\infty} d\mu \left. \frac{d}{d\nu} K_{\nu,\mu}(\varphi) \right|_{\nu=0} \\ &= -\frac{1}{2} \int_0^{+\infty} \frac{d\mu}{\mu} (\varphi, i ([Q^2, \chi_\mu(P)] + [Q^2, \chi_{-\mu}(P)]) \varphi). \end{aligned} \quad (4.22)$$

Due to Lemma 2.2.(a) we have

$$\begin{aligned}
& \int_0^{+\infty} \frac{d\mu}{\mu} (\varphi, [Q^2, \chi_\mu(P)]\varphi) \\
&= \lim_{\varepsilon \searrow 0} \int_\varepsilon^{+\infty} \frac{d\mu}{\mu} [(Q^2\varphi, \chi_\mu(P)\varphi) - (\chi_\mu(P)\varphi, Q^2\varphi)] \\
&= \lim_{\varepsilon \searrow 0} [(Q^2\varphi, (\int_\varepsilon^{+\infty} \frac{d\mu}{\mu} \chi_\mu(P) + \ln \varepsilon)\varphi) - ((\int_\varepsilon^{+\infty} \frac{d\mu}{\mu} \chi_\mu(P) + \ln \varepsilon)\varphi, Q^2\varphi)] \\
&= (\varphi, [Q^2, R_\Sigma(P)]\varphi).
\end{aligned}$$

This together with (4.22) lead to the desired equality, that is

$$K_\infty(\varphi) = -\frac{1}{2} (\varphi, i[Q^2, R_\Sigma(P) + R_\Sigma(-P)]\varphi) = -(\varphi, i[Q^2, G_\Sigma(P)]\varphi).$$

(iii) To apply Lebesgue's dominated convergence theorem to (4.21) we need to bound $K_{\nu,\mu}(\varphi)$ uniformly in ν by a function in $L^1(\mathbb{R}_+, d\mu)$. We do this separately for $\mu \leq 1$ and for $\mu \geq 1$.

We begin with the case $\mu \leq 1$. Write $K_{\nu,\mu}(\varphi)$ as

$$K_{\nu,\mu}(\varphi) = F_{\nu,\mu}(\varphi) - F_{\nu,-\mu}(\varphi),$$

where

$$\begin{aligned}
F_{\nu,\mu}(\varphi) &= \frac{1}{\nu\mu^2} [(Z_{\nu\mu}\varphi, \chi_\mu(P)Z_{\nu\mu}\varphi) - (\varphi, \chi_\mu(P)\varphi)] \\
&= \frac{1}{\mu} \left(\left(\frac{Z_{\nu\mu}-\mathbb{1}}{\nu\mu} \right) \varphi, \chi_\mu(P)Z_{\nu\mu}\varphi \right) + \frac{1}{\mu} (\chi_\mu(P)\varphi, \left(\frac{Z_{\nu\mu}-\mathbb{1}}{\nu\mu} \right) \varphi).
\end{aligned}$$

Due to the spectral theorem, we have

$$\left\| \left(\frac{Z_{\pm\nu\mu}-\mathbb{1}}{\nu\mu} \right) \varphi \right\| \leq \text{Const.} \|\langle Q \rangle^2 \varphi\|. \quad (4.23)$$

Let $0 < \ell < \frac{1}{2}$, then $|P|^{-\ell} \langle Q \rangle^{-2}$ is bounded (after conjugation by a Fourier transform this follows from the fact that $|Q|^{-\ell}$ is P^2 -bounded [1, Prop. 2.28]). Since Σ is bounded, we have

$$|\mu^{-1} \xi|^\ell |\chi_{\pm\mu}(\xi)| \leq \text{Const.}$$

Therefore

$$\begin{aligned}
\mu^{-1} \|\chi_{\pm\mu}(P)\varphi\| &= \mu^{\ell-1} \|\mu^{-1} P^\ell \chi_{\pm\mu}(P) |P|^{-\ell} \langle Q \rangle^{-2} \langle Q \rangle^2 \varphi\| \\
&\leq \text{Const.} \mu^{\ell-1} \|\langle Q \rangle^2 \varphi\|,
\end{aligned} \quad (4.24)$$

and

$$\mu^{-1} \|\chi_{\pm\mu}(P)Z_{\pm\nu\mu}\varphi\| \leq \text{Const.} \mu^{\ell-1} \|\langle Q \rangle^2 \varphi\|. \quad (4.25)$$

From (4.23), (4.24) and (4.25) we get the estimates

$$|F_{\nu,\pm\mu}(\varphi)| \leq \text{Const.} \mu^{\ell-1} \|\langle Q \rangle^2 \varphi\|^2.$$

Thus we have

$$|K_{\nu,\mu}(\varphi)| \leq \text{Const.} \mu^{\ell-1} \|\langle Q \rangle^2 \varphi\|^2, \quad (4.26)$$

which shows that $K_{\nu,\mu}(\varphi)$ is bounded uniformly in ν by a function in $L^1([0, 1], d\mu)$.

We consider now the case $\mu \geq 1$. Since $\varphi = \mathbb{1}_J(H_0)\varphi$ for some compact set J , there exists $\mu_0 \geq 0$ such that

$$(\varphi, \chi_\mu(P)\varphi) = (\varphi, \chi_{-\mu}(P)\varphi) = (\varphi, \varphi), \quad \forall \mu \geq \mu_0.$$

Hence for $\mu \geq \mu_0$, we have

$$\begin{aligned} |K_{\nu,\mu}(\varphi)| &= \frac{1}{\nu\mu^2} |(\varphi, Z_{-\nu\mu}\chi_\mu(P)Z_{\nu\mu}\varphi) - (\varphi, Z_{\nu\mu}\chi_{-\mu}(P)Z_{-\nu\mu}\varphi)| \\ &= \frac{1}{\nu\mu^2} |(\|\chi_\mu(P)Z_{\nu\mu}\varphi\|^2 - \|Z_{\nu\mu}\varphi\|^2) - (\|\chi_{-\mu}(P)Z_{-\nu\mu}\varphi\|^2 - \|Z_{-\nu\mu}\varphi\|^2)| \\ &\leq \frac{2}{\nu\mu^2} \|\varphi\| \left[\|(\chi_\mu(P) - \mathbb{1})Z_{\nu\mu}\varphi\| + \|(\chi_{-\mu}(P) - \mathbb{1})Z_{-\nu\mu}\varphi\| \right], \end{aligned} \quad (4.27)$$

where we have used (4.11) in the last step. To bound the r.h.s. of (4.27) we will use the following identity, which is an easy consequence of (4.19):

$$[F(P + sQ) - F(P)]\varphi = \frac{1}{2} \int_0^s d\tau [2(\nabla F)(P + \tau Q) \cdot Q - i(\Delta F)(P + \tau Q)]\varphi, \quad (4.28)$$

where F is any bounded function in $C^\infty(\mathbb{R}^d)$ with bounded derivatives.

Let $F \in C^\infty(\mathbb{R}^d)$ with $F \equiv 1$ near infinity, $F \equiv 0$ near 0 be such that

$$F(Q)[\chi(Q) - \mathbb{1}] = \chi(Q) - \mathbb{1}.$$

Then we have

$$\|[\chi_{\pm\mu}(P) - \mathbb{1}]Z_{\pm\nu\mu}\varphi\| \leq \|F(\frac{\pm P}{\mu})Z_{\pm\nu\mu}\varphi\| = \|F(\frac{\pm P}{\mu} + \nu Q)\varphi\| \quad (4.29)$$

due to (4.19). From (4.29) and the fact that

$$F(\frac{\pm P}{\mu})\varphi = 0, \quad \forall \mu \geq \mu_0,$$

we get for $\mu \geq \mu_0$

$$|K_{\nu,\mu}(\varphi)| \leq \frac{2}{\nu\mu^2} \|\varphi\| \left\{ \|[F(\frac{P}{\mu} + \nu Q) - F(\frac{P}{\mu})]\varphi\| + \|[F(\frac{-P}{\mu} + \nu Q) - F(\frac{-P}{\mu})]\varphi\| \right\}$$

Moreover one has

$$[F(\frac{\pm P}{\mu} + \nu Q) - F(\frac{\pm P}{\mu})]\varphi = \frac{1}{2} \int_0^\nu d\tau [2(\nabla F)(\frac{\pm P}{\mu} + \tau Q) \cdot Q - i(\Delta F)(\frac{\pm P}{\mu} + \tau Q)]\varphi$$

due to (4.28). Therefore we have for $\mu \geq \mu_0$

$$|K_{\nu,\mu}(\varphi)| \leq \text{Const.} \mu^{-2} \|\langle Q \rangle \varphi\|^2. \quad (4.30)$$

The combination of (4.26) and (4.30) shows that $K_{\nu,\mu}(\varphi)$ is bounded uniformly in ν by a function in $L^1([1, +\infty[, d\mu)$. \square

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